

On the Einstein—Cartan Equations. II.

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Summary. It is shown that by covariant exterior differentiation of the equations of the generalized theory of gravity one arrives at a system of relations which hold as an algebraic consequence of the field equations themselves. The Dirac equation in a Riemann-Cartan space-time is written in a manner adapted to the calculus of exterior forms.

1. The differential identities. The curvature and torsion forms of any affinely connected space satisfy the Bianchi identities*)

$$D\Omega_l^k = 0 \quad \text{and} \quad D\Theta^k = \Omega_l^k \wedge \theta^l.$$

These identities may be used to evaluate the covariant exterior derivatives of the 3-forms e_j and c_{kl} appearing in the Einstein—Cartan equations of the generalized, metric theory of gravitation. It follows directly from (I.11) and (I.12) that

$$(1) \quad De_j = \frac{1}{2} \eta_{jklm} \Theta^m \wedge \Omega^{kl},$$

$$(2) \quad Dc_{kl} = \eta_{lj} \wedge \Omega_k^j - \eta_{kj} \wedge \Omega_l^j.$$

Eq. (1) is the generalization, to the Riemann-Cartan space, of the 'contracted Bianchi identity' $De_j = 0$ which plays an important role in Einstein's theory of gravitation. The right-hand sides of Eqs (1) and (2) may be rearranged to give the formulae

$$(3) \quad De_j = Q_j^k \wedge e_k + \frac{1}{2} R^{kl}{}_j \wedge c_{kl},$$

$$(4) \quad Dc_{kl} = \theta_k \wedge e_l - \theta_l \wedge e_k,$$

where

$$\frac{1}{2} \theta^j \wedge Q_j^k = \Theta^k \quad \text{and} \quad \frac{1}{2} \theta^j \wedge R^{kl}{}_j = \Omega^{kl}.$$

Let σ be a representation of the Lorentz group in \mathbf{R}^N and let (σ_{Ak}^{Bl}) be the matrix corresponding to the derived homomorphism of Lie algebras. The matrix $\sigma_{kl} =$

*) In this note, similar assumptions are made, and the same notation is used, as in Part I [1]. Equations appearing there are referred to in the style (I.n). The frames (θ^k) are assumed to be orthonormal so that $g_{kl} = 0$ for $k \neq l$ and $g_{11} = g_{22} = g_{33} = -g_{44} = -1$.

