

On the Einstein—Cartan Equations. II.

by

A. TRAUTMAN

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**Summary.** It is shown that by covariant exterior differentiation of the equations of the generalized theory of gravity one arrives at a system of relations which hold as an algebraic consequence of the field equations themselves. The Dirac equation in a Riemann-Cartan space-time is written in a manner adapted to the calculus of exterior forms.

**1. The differential identities.** The curvature and torsion forms of any affinely connected space satisfy the Bianchi identities\*)

$$D\Omega_l^k = 0 \quad \text{and} \quad D\theta^k = \Omega_l^k \wedge \theta^l.$$

These identities may be used to evaluate the covariant exterior derivatives of the 3-forms  $e_j$  and  $c_{kl}$  appearing in the Einstein—Cartan equations of the generalized, metric theory of gravitation. It follows directly from (I.11) and (I.12) that

$$(1) \quad De_j = \frac{1}{2} \eta_{jklm} \theta^m \wedge \Omega^{kl},$$

$$(2) \quad Dc_{kl} = \eta_{lj} \wedge \Omega_k^j - \eta_{kj} \wedge \Omega_l^j.$$

Eq. (1) is the generalization, to the Riemann-Cartan space, of the 'contracted Bianchi identity'  $De_j = 0$  which plays an important role in Einstein's theory of gravitation. The right-hand sides of Eqs (1) and (2) may be rearranged to give the formulae

$$(3) \quad De_j = Q_j^k \wedge e_k + \frac{1}{2} R^{kl}{}_j \wedge c_{kl},$$

$$(4) \quad Dc_{kl} = \theta_k \wedge e_l - \theta_l \wedge e_k,$$

where

$$\frac{1}{2} \theta^j \wedge Q_j^k = \theta^k \quad \text{and} \quad \frac{1}{2} \theta^j \wedge R^{kl}{}_j = \Omega^{kl}.$$

Let  $\sigma$  be a representation of the Lorentz group in  $\mathbf{R}^N$  and let  $(\sigma_{Ak}^{Bl})$  be the matrix corresponding to the derived homomorphism of Lie algebras. The matrix  $\sigma_{kl} =$

\*) In this note, similar assumptions are made, and the same notation is used, as in Part I [1]. Equations appearing there are referred to in the style (I.n). The frames  $(\theta^k)$  are assumed to be orthonormal so that  $g_{kl} = 0$  for  $k \neq l$  and  $g_{11} = g_{22} = g_{33} = -g_{44} = -1$ .

$= (\sigma_{Akl}^B, \sigma_{Akl}^B = g_{jl} \sigma_{Ak}^{Bj}$ , is skew,  $\sigma_{kl} + \sigma_{lk} = 0$ . In the metric theory, the forms  $\omega^{kl} = g^{jl} \omega_j^k$  of the connection, referred to a field  $(\theta^k)$  of orthonormal frames, are also skew in the pair  $(k, l)$ . The covariant derivative of a tensor-valued 0-form of type  $\sigma$ ,  $\varphi = (\varphi_A)$ , is given by

$$D\varphi = d\varphi + \omega^{kl} \sigma_{kl} \varphi = \varphi_j \theta^j.$$

The indices  $A, B = 1, \dots, N$  have been suppressed in this formula, which should be interpreted similarly as in spinor calculus.

Let the lagrangian form corresponding to  $\varphi$  be  $L = \eta A(\varphi, \varphi_j)$ , where  $A$  is a function on  $\mathbf{R}^N \times \mathcal{L}(\mathbf{R}^4, \mathbf{R}^N)$ . In this formulation, where everything is referred to orthonormal frames, there is no possibility for considering changes of  $g_{ij}$ . A general variation of  $L$  is

$$\delta L = t_j \wedge \delta \theta^j + \frac{1}{2} s_{kl} \wedge \delta \omega^{kl} + L^A \delta \varphi_A + \text{an exact form,}$$

where

$$(5) \quad t_j = \left( \frac{\partial A}{\partial \varphi_k} \varphi_j - A \delta_j^k \right) \eta_k, \quad s_{kl} = -2 \frac{\partial A}{\partial \varphi_j} \sigma_{kl} \varphi \eta_j,$$

and

$$L^A = \eta \frac{\partial A}{\partial \varphi_A} - D \left( \eta_j \frac{\partial A}{\partial \varphi_{Aj}} \right).$$

A mere Lorentz rotation of the frames leaves both  $L$  and  $A$  invariant. This invariance results in the equation

$$(6) \quad Ds_{kl} = \theta_k \wedge t_l - \theta_l \wedge t_k \quad \text{mod } L^A = 0$$

proving the symmetry of the tensor  $T_{kl}$  which should now be considered as *defined* by Eq. (I.6). The covariant exterior derivative of  $t_j$  may be evaluated from Eq. (5)

$$(7) \quad Dt_j = Q_j^k \wedge t_k + \frac{1}{2} R^{kl} \wedge s_{kl} \quad \text{mod } L^A = 0.$$

A simple comparison of Eqs. (I.8) with (3), (4), (6) and (7) leads to

**THEOREM.** *The equations resulting by covariant exterior derivation of the Einstein-Cartan equations are algebraic consequences of the Einstein-Cartan equations themselves and of the equations of motion of matter.*

The implications of this result should be compared with, and distinguished from, the point of view of Cartan who required that  $Dt_j$  be zero ([2], pp. 21–23).

**2. The Dirac equation in a space-time with torsion.** To illustrate the general formalism developed in [1] and in the preceding section, we give here a brief account of the Dirac equation appropriate to a Riemann-Cartan space-time. A detailed and different treatment of this subject may be found elsewhere [3].

Let  $X$  be a four-dimensional differential manifold endowed with a metric tensor  $g$  of hyperbolic signature. A simple-minded, physicist's approach to Dirac spinors on  $X$  may be summarized as follows. The Dirac matrices  $\gamma_k \in \mathcal{L}(C^4)$  satisfy

$$\gamma_k \gamma_l + \gamma_l \gamma_k = 2g_{kl}$$

and  $\beta = \beta^+$  is a matrix such that

$$\gamma_k^+ = \beta \gamma_k \beta^{-1},$$

where cross denotes hermitean conjugation. The six spin matrices

$$\sigma_{kl} = \frac{1}{8} (\gamma_k \gamma_l - \gamma_l \gamma_k)$$

satisfy

$$(8) \quad \sigma_{kl} \gamma_j - \gamma_j \sigma_{kl} = \frac{1}{2} (g_{jl} \gamma_k - g_{jk} \gamma_l),$$

$$(9) \quad \sigma_{kl} \gamma_j + \gamma_j \sigma_{kl} = \frac{1}{2} \eta_{jklm} \gamma_5 \gamma^m,$$

where  $\gamma_5 = \gamma_1 \gamma_2 \gamma_3 \gamma_4$ ,  $\gamma_k = g_{kl} \gamma^l$ ,  $\eta_{1234} = 1$ , etc. Space-time is assumed to have a global field of orthonormal frames ( $\theta^k$ ). One considers only those fields of orthonormal frames ( $\theta'^k$ ) which differ 'little' from a given field ( $\theta^k$ ),

$$\theta'^k = \theta^k + \delta\theta^k = \theta^k - a_k^i \theta^i \quad \text{and} \quad a_{kl} + a_{lk} = 0.$$

A spinor field on  $X$  is a law which associates to each of the fields of frames a map from  $X$  to  $\mathbb{C}^4$  in such a way that if  $\psi: X \rightarrow \mathbb{C}^4$  corresponds to ( $\theta^k$ ) and  $\psi' = \psi + \delta\psi$  corresponds to ( $\theta'^k$ ), then

$$\delta\psi = -a^{kl} \sigma_{kl} \psi.$$

Similarly, for the contragredient spinor  $\bar{\psi} = \psi^+ \beta$ ,

$$\delta\bar{\psi} = \bar{\psi} \sigma_{kl} a^{kl}.$$

Let  $X$  be endowed, in addition to the metric tensor, with a metric affine connection described by the collection ( $\omega_{kl}$ ) of 1-forms. The covariant derivative  $D\psi$  of a spinor field  $\psi$  is a spinor-valued 1-form

$$D\psi = d\psi + \omega^{kl} \sigma_{kl} \psi.$$

The corresponding formula for the derivative of the contragredient spinor is

$$D\bar{\psi} = d\bar{\psi} - \bar{\psi} \sigma_{kl} \omega^{kl}.$$

The covariant exterior derivative of  $\iota = \gamma_k \eta^k$  (the dual of  $\gamma = \gamma_k \theta^k$ ),

$$D\iota = \iota \wedge Q_k^k$$

depends on the trace of the torsion tensor. If  $*D\psi$  denotes the dual of  $D\psi$ , then

$$(10) \quad -\iota \wedge D\psi = \gamma \wedge *D\psi = \eta \gamma^k \nabla_k \psi.$$

The lagrangian form corresponding to a Dirac particle of mass  $m$  is

$$L = \frac{1}{2} i (\bar{\psi} \iota \wedge D\psi + D\bar{\psi} \wedge \iota \psi) + m \eta \bar{\psi} \psi.$$

The Dirac equation obtained by varying the action integral with respect to  $\bar{\psi}$  or  $\psi$ ,

$$(11) \quad \iota \wedge D\psi - D(\iota \psi) = 2im\eta \psi,$$

$$(12) \quad (D\bar{\psi}) \wedge \iota + D(\bar{\psi} \iota) = 2im\eta \bar{\psi},$$

implies the conservation law  $dj=0$ , where  $j=\bar{\psi}\gamma^j\psi$  is the current. By varying  $L$  with respect to  $\theta^j$  and  $\omega^{kl}$ , and making use of Eqs (8)–(12), one obtains

$$(13) \quad t_j = \frac{1}{2} i (\bar{\psi}\gamma_j *D\psi - *D\bar{\psi}\gamma_j\psi)$$

and

$$(14) \quad s_{kl} = i\bar{\psi} (\sigma_{kl} + \sigma_{kl}^t) \psi = -\frac{1}{2} i\bar{\psi}\gamma_5 \gamma^j\psi \wedge \theta_k \wedge \theta_l.$$

INSTITUTE OF THEORETICAL PHYSICS, UNIVERSITY, WARSAW  
(INSTYTUT FIZYKI TEORETYCZNEJ, UNIWERSYTET, WARSZAWA)

#### REFERENCES

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A. Траутман, **Об уравнениях Эйнштейна — Картана. II.**

**Содержание.** В настоящей работе показано, что при использовании ковариантного внешнего дифференцирования уравнений обобщенной теории гравитации, получаем систему соотношений, которая является алгебраическим следствием самых уравнений поля. Уравнение Дирака в пространстве — времени Римана-Картана записано при использовании формализма внешних форм.